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ASTRONOMICAL TESTS FOR
EXTENDED GRAVITY: POSSIBLE
CONSTRAINTS ON $f(R)$ WITH
VANISHING COSMOLOGICAL
CONSTANT

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Dark matter, dark energy

new physics

new geometry



Dark matter, dark energy

New physics

New geometry



scalar-tensor gravity

Dark matter, dark energy

New physics

New geometry



Scalar-tensor gravity

$f(R)$ gravity

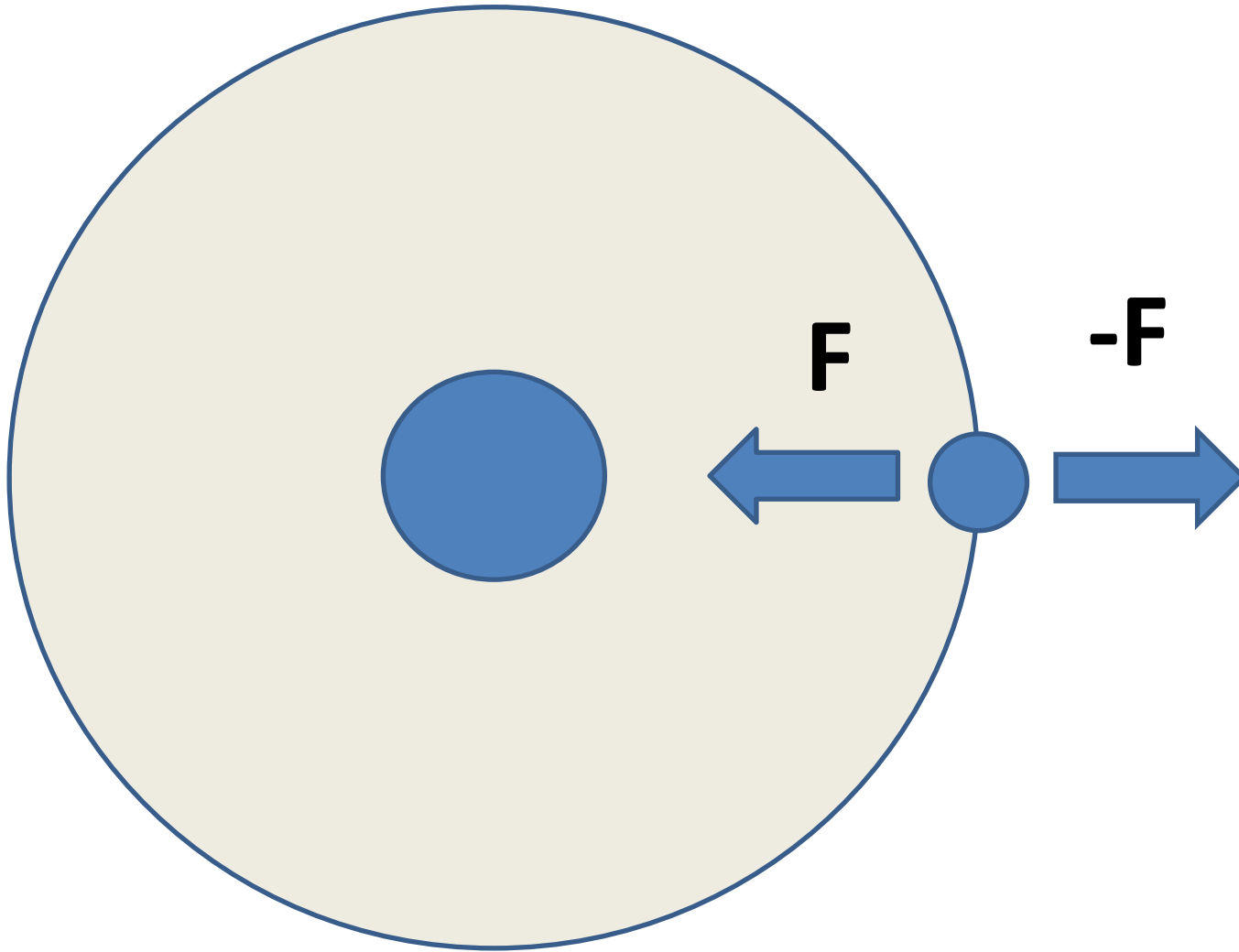
Dark energy + f(R) gravity



Starobinsky model

$$f(R) = R + \lambda R_0 \left(\left(1 + \frac{R^2}{R_0^2} \right)^{-n} - 1 \right)$$

Method: cut-off radius calculation



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- Starobinsky model
- Λ CDM

Mass spectra:

$10^{11} M_{\text{Sun}}$ (Milky Way) – $10^{15} M_{\text{Sun}}$ (Galaxy clusters)

- **Metrics**

$$ds^2 = e^A c^2 dt^2 - e^{-A} dr^2 - r^2 d\Omega.$$

- **Field equations***

$$f'(R)R_{ii} - f(R)\frac{g_{ii}}{2} - (\nabla_i^2 - g_{ii}\square) f'(R) = 0.$$

*Thomas P. Sotiriou and Valerio Faraoni, Rev. Mod. Phys. 82, p.451 (2010)

- **Field equations**

$$\left\{ \begin{array}{l} \frac{dA}{dr} = a, \\ \frac{da}{dr} = e^{-A} \left(R + \frac{2}{r^2} \right) - a^2 - \frac{4a}{r} - \frac{2}{r^2} \\ \frac{dR}{dr} = \frac{f'(R) \cdot R_{00}/g_{00} - f(R)/2}{(e^A(a + 2/r) + a/2)f''(R)} \end{array} \right.$$

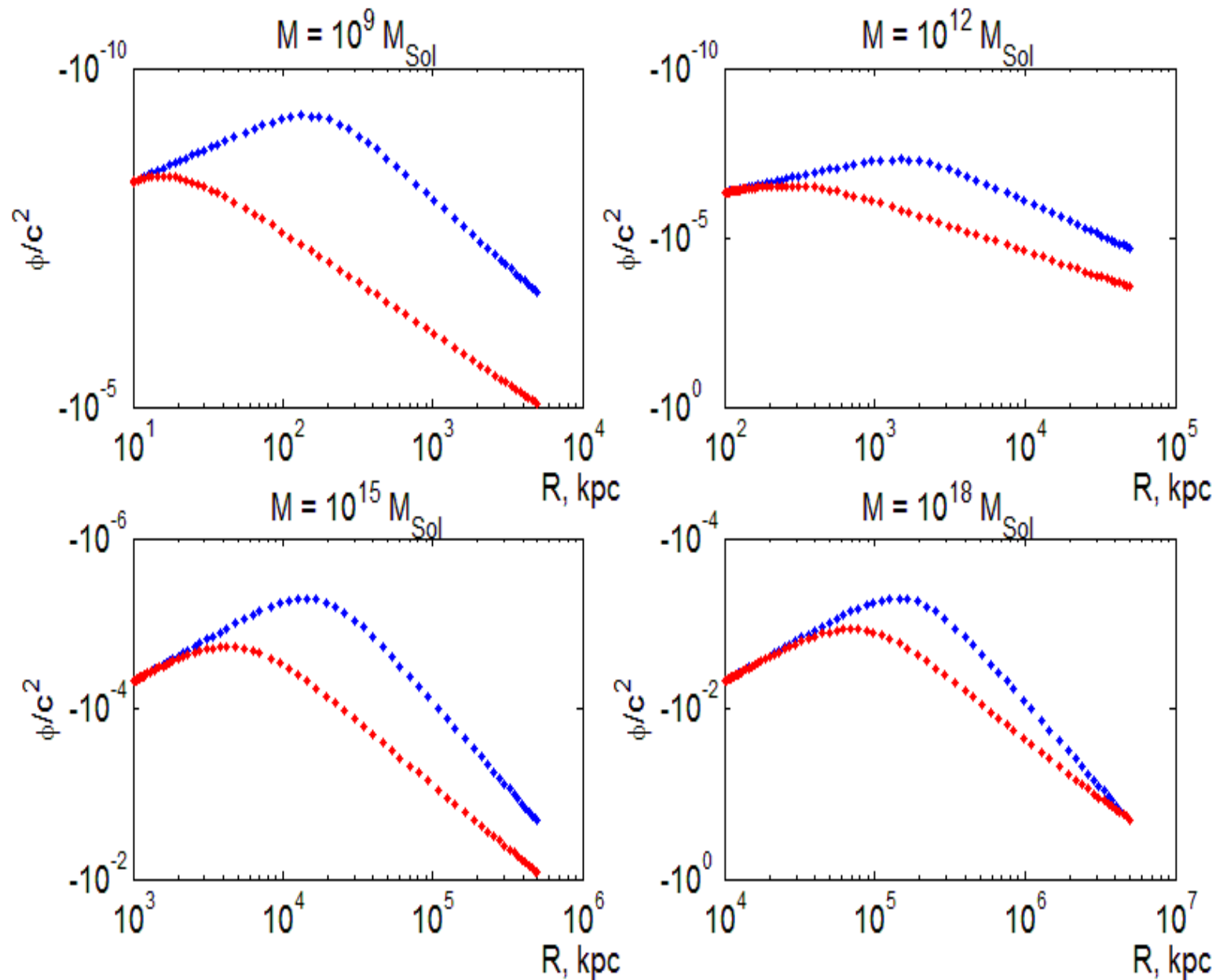
- **Gravitational potential**

$$\phi = \frac{c^2}{2}(g_{00} - 1) = \frac{c^2}{2}(e^A - 1);$$

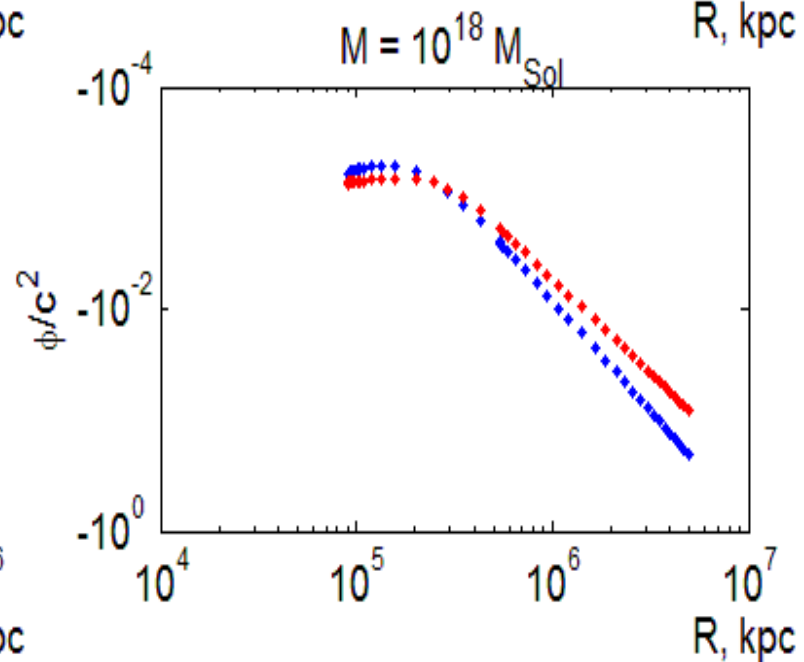
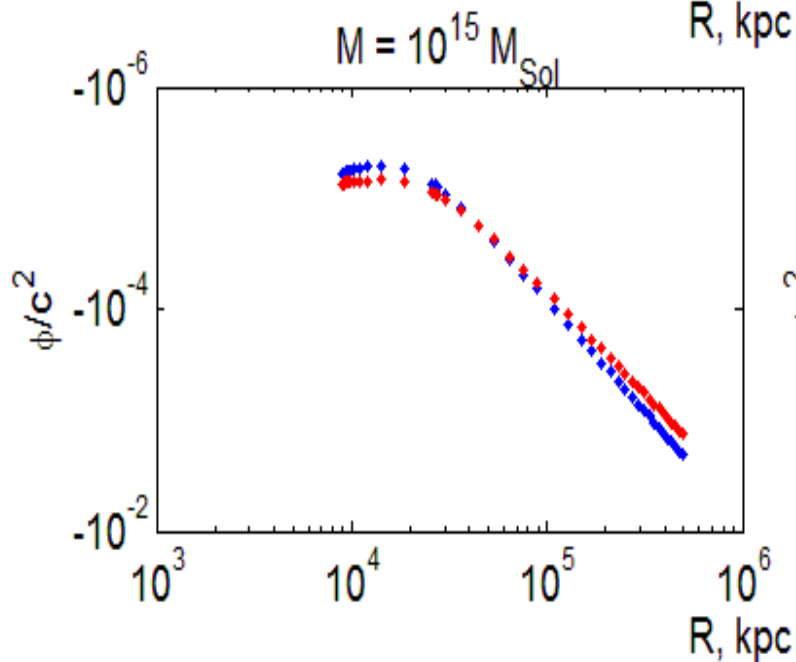
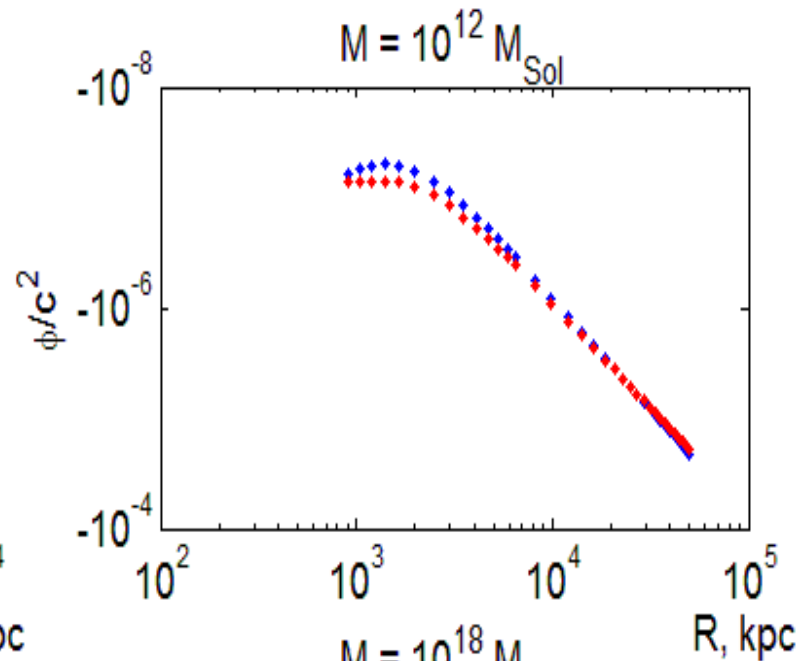
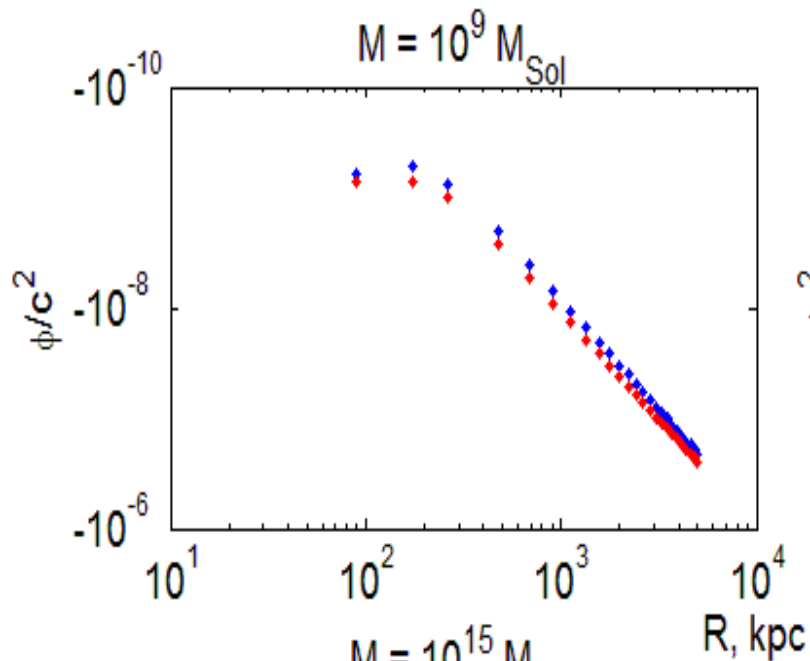
- **Dimensionless gravitational potential**

$$\phi/c^2 = \frac{e^A - 1}{2};$$

ϕ/c^2 , case $n=1$



ϕ/c^2 , case $n=9$



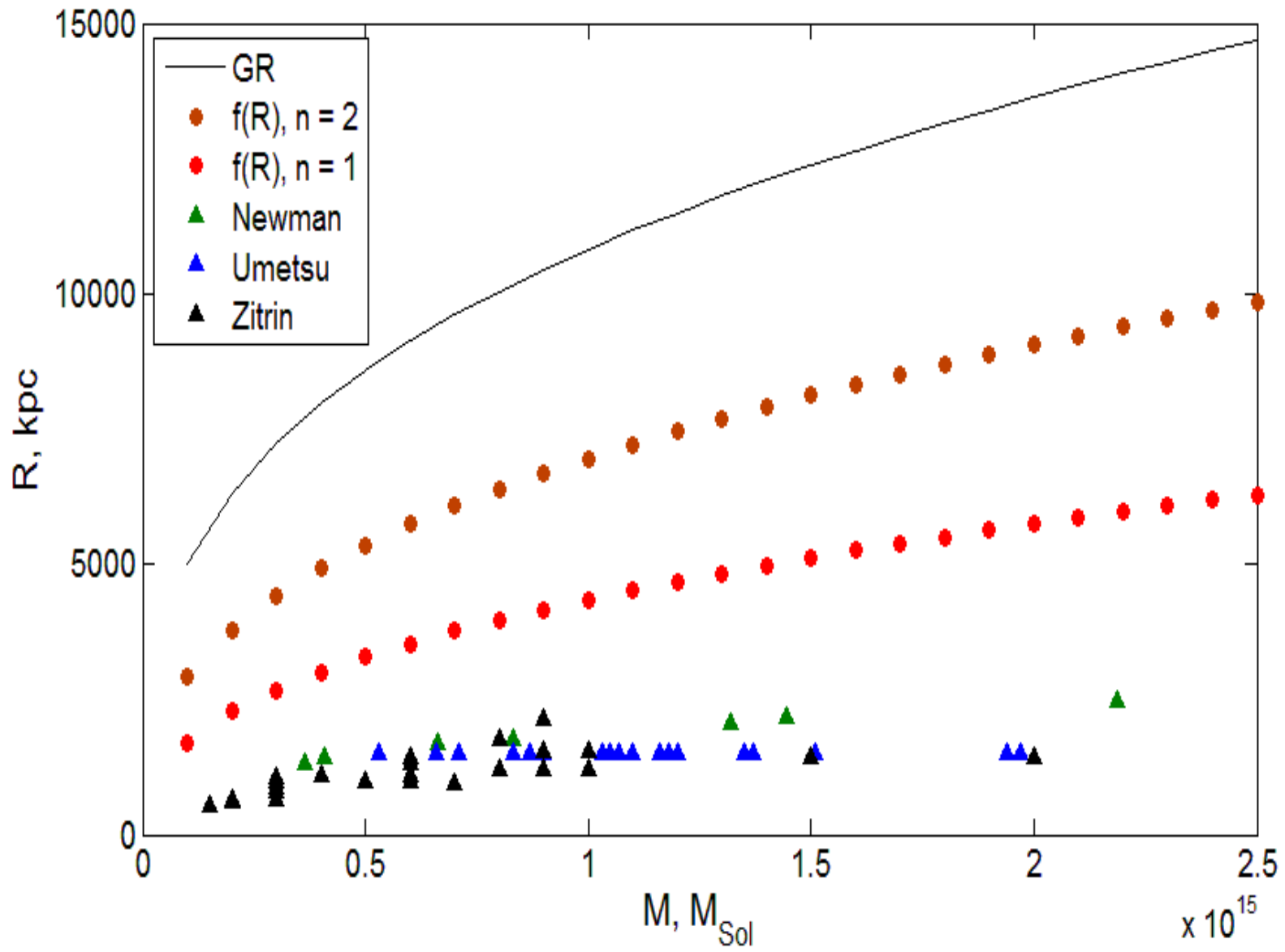
(Numerically
obtained)
coefficients
for $n=1..9$

$\log_{10}(R_{\text{cut-off}})$

$= \alpha \log_{10}(M) + \beta$

n	α	β
0	0.3333	-0.966
1	0.4084±0.0013	-2.495±0.017
2	0.3807±0.0004	-1.869±0.005
3	0.3683±0.0001	-1.578±0.001
4	0.3617±0.0001	-1.412±0.002
5	0.3577±0.0002	-1.305±0.003
6	0.3552±0.0003	-1.231±0.004
7	0.3528±0.0005	-1.169±0.007
8	0.3486±0.0015	-1.091±0.019
9	0.3428±0.0019	-0.993±0.024

Cut-off radius: theory & observations



Conclusions

The dependence of cut-off radius versus mass for Starobinsky model with vanishing cosmological constant has no visible now difference from GR one in the wide range of masses



The current accuracy value does not provide additional restrictions and constraints on Starobinsky model

THANK YOU FOR YOUR
ATTENTION !

