

Recent developments in Neutrino Physics

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Plan

Oscillation sector

- Standard 3-neutrino oscillation
- Anomalies in neutrino data

Flavor sector

- Problem of neutrino masses and mixings:
the role of family symmetries

Introduction

- New data from neutrino oscillation experiments have given precise results on mixing parameters
- Possible leptonic CP violation ($\leq 5 \sigma$) (T2K, NoVA...)

- **However:**

not a unique extension of the Standard Model that allows to explain:

- origin of masses and mixing angles
- differences with respect to the quark sector

3- ν state formalism

- Neutrinos can be described in terms of mass or weak eigenstates



$$|\nu_\alpha\rangle = \sum_{i=1}^3 U_{\alpha i} |\nu_i\rangle$$

neutrino matrix
matrix

- Simple time evolutions of the vector $\nu(t) = (\nu_e(t), \nu_\mu(t), \nu_\tau(t))$:

$$i \frac{d}{dt} |\nu(t)\rangle = H |\nu(t)\rangle$$

$$H = \frac{1}{2E_\nu} U \text{Diag}[0, m_2^2 - m_1^2, m_3^2 - m_1^2] U^\dagger$$

there exist a
probability of a change
of the neutrino flavour

Theory of neutrino oscillation

- Flavour changing transitions

$$P(\nu_\alpha \rightarrow \nu_\beta) = |\langle \nu_\beta | \nu_\alpha(t) \rangle|^2 = \left| \sum_j U_{\beta j} e^{\frac{-i m_j^2 L}{2 E_\nu}} U_{\alpha j}^* \right|^2$$

$\alpha = \beta \rightarrow$ disappearance

$\alpha \neq \beta \rightarrow$ appearance

- In the case of two neutrinos only:

$$P(\nu_e \rightarrow \nu_\mu) = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4 E_\nu} \right)$$

distance source-detector

mixing angle

$$U = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$$

U = unitary matrix

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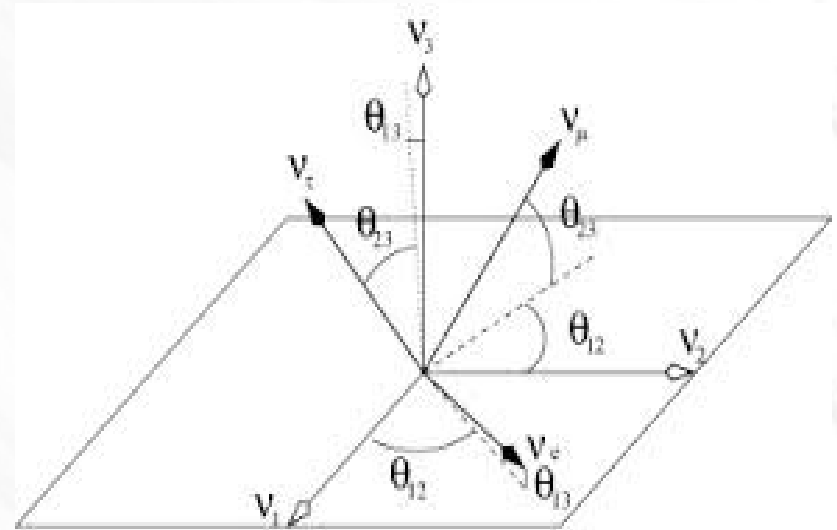
Бруно Понтекорво

Theory of neutrino oscillation

- The neutrino mixing matrix depends on 4 real parameters

$$U_{PMNS} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\text{atmospheric mixing}} \times \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13} e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{-i\delta} & 0 & c_{13} \end{pmatrix}}_{\text{reactor mixing}} \times \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{solar mixing}}$$

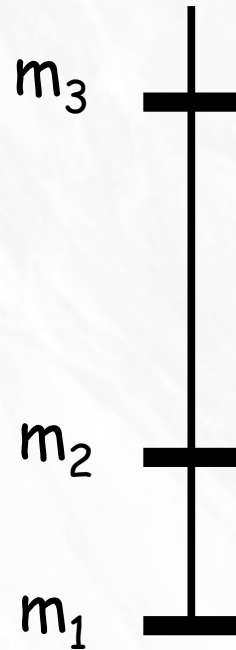
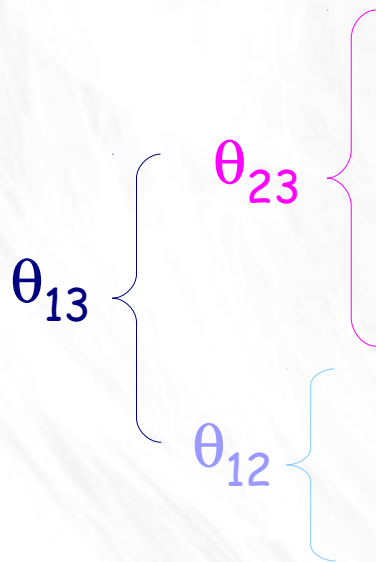
- A more complicated mismatch between the ν bases



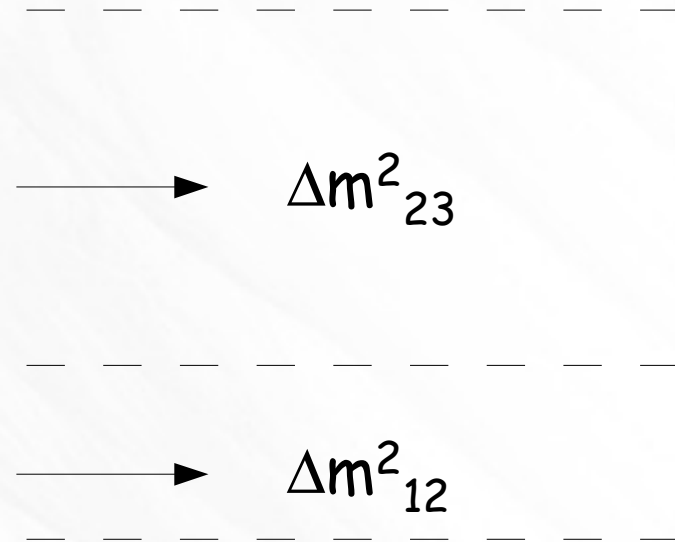
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Experiments measure...

Mixing angles



Mass differences



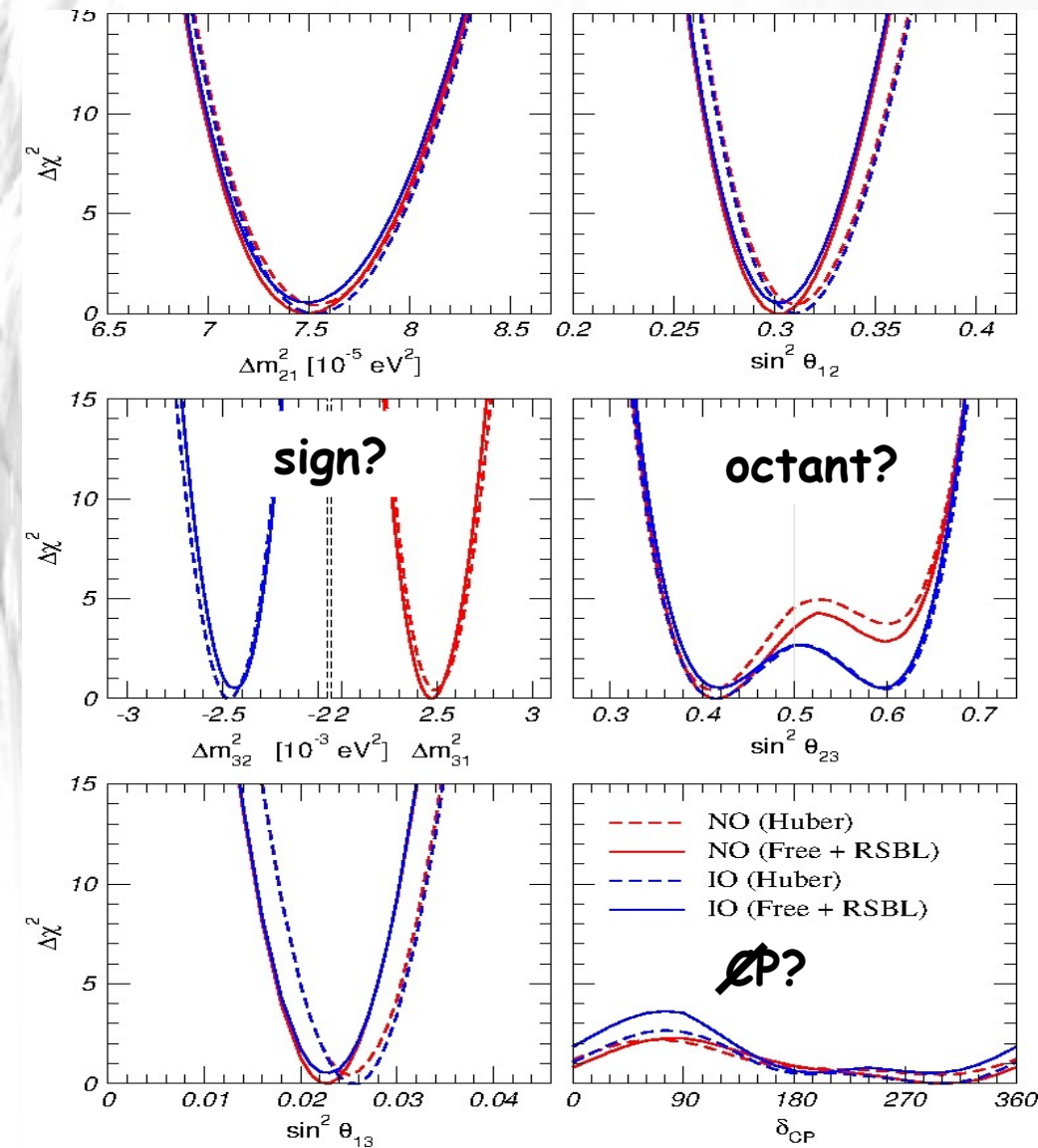
- unknowns: leptonic CP violation and the ordering of the mass eigenstates

Global fit

Gonzalez-Garcia et al. JHEP1212,(2012)123

Parameter	Result
θ_{12}	$33.36^{+0.81}_{-0.78}$
θ_{13}	$8.66^{+0.44}_{-0.46}$
θ_{23}	$40.0^{+2.1}_{-1.5}$
δ	300^{+66}_{-138}
$\Delta m^2_{23} (10^{-3} \text{ eV}^2)$	$2.47^{+0.07}_{-0.07}$
$\Delta m^2_{12} (10^{-5} \text{ eV}^2)$	$7.50^{+0.18}_{-0.19}$

- Masses @3%
- Angles between 5% and 10%

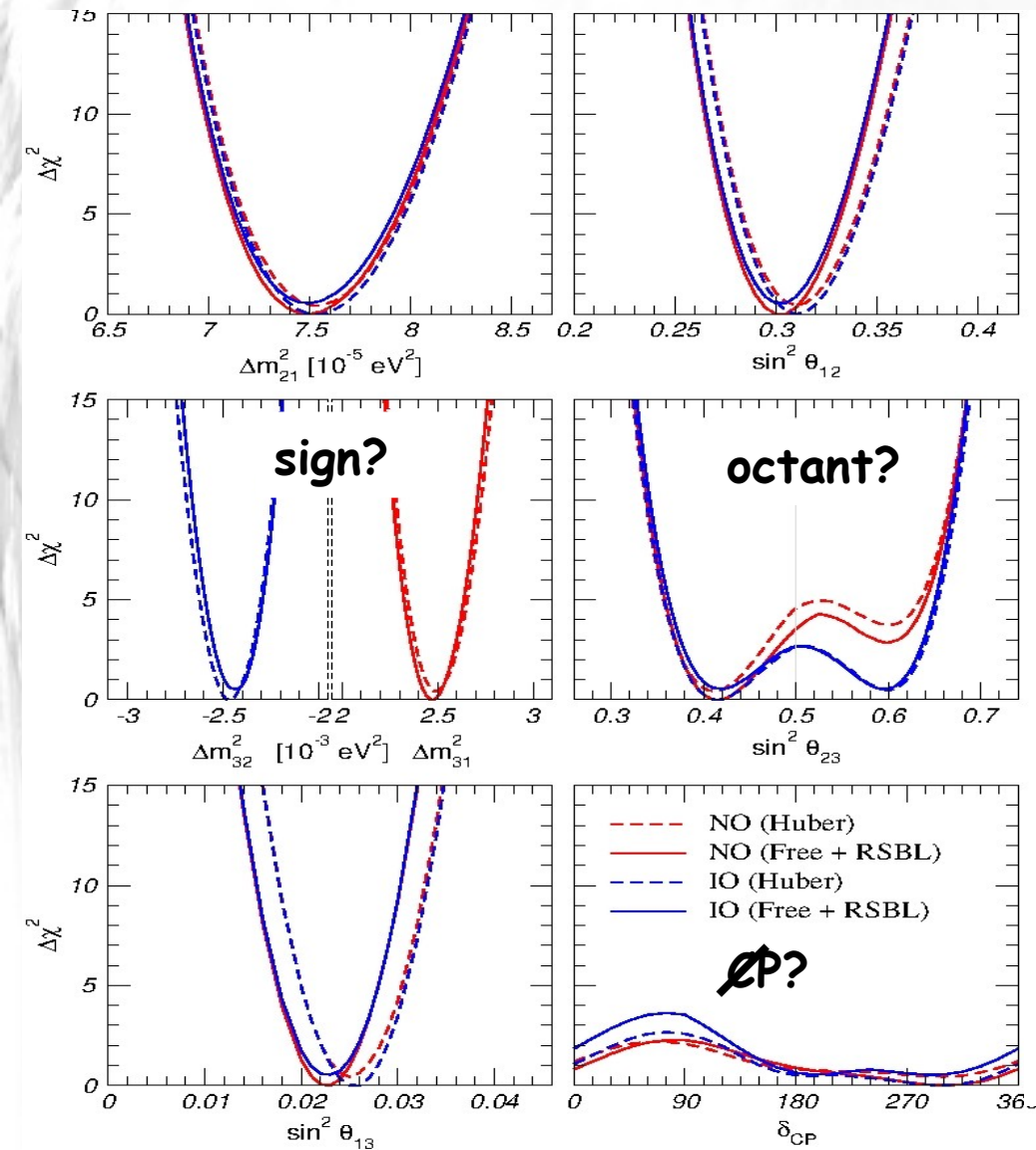


Global fit

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$$|U| = \begin{pmatrix} 0.795 \rightarrow 0.846 & 0.513 \rightarrow 0.585 & 0.126 \rightarrow 0.178 \\ 0.205 \rightarrow 0.543 & 0.416 \rightarrow 0.730 & 0.579 \rightarrow 0.808 \\ 0.215 \rightarrow 0.548 & 0.409 \rightarrow 0.725 & 0.567 \rightarrow 0.800 \end{pmatrix}$$



Neutrino oscillation anomalies

- LSND

evidence for oscillations $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ with $L/E \sim 1 \text{ km/GeV}$ (ν_e appearance)

- Anomalies in Gallium experiments (SAGE & GALLEX)

they measured an electron neutrino flux from the Sun smaller than expected (ν_e disappearance)

- Anomalies due to new computations of reactor neutrino fluxes

fluxes from reactor neutrinos are $\sim 3.5\%$ larger than in the past \rightarrow
experiments with $L \leq 100 \text{ m}$ show deficit of neutrinos (ν_e disappearance-
Bugey, Rovno...)

In addition there are *null results*: ν_μ disappearance (CDHS, SK, MINOS) e ν_e appearance (KARMEN, NOMAD, ICARUS, OPERA) which gave **no signal**

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Neutrino oscillation anomalies

Joachim Kopp

August 21, Aspen

- LSND

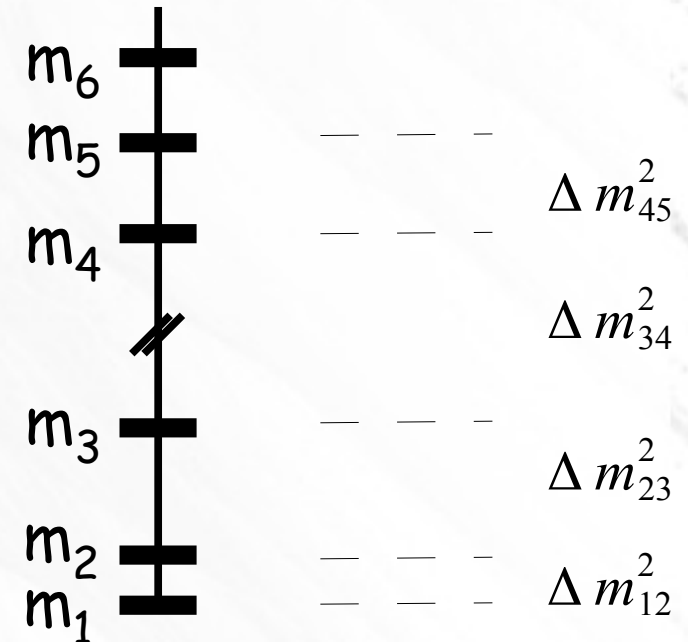
evidence for oscillations

$\bar{\nu}_e \rightarrow \bar{\nu}_\mu$ con $L/E \sim 1$
km/GeV

- MiniBooNE

no significant excess of ν_e o

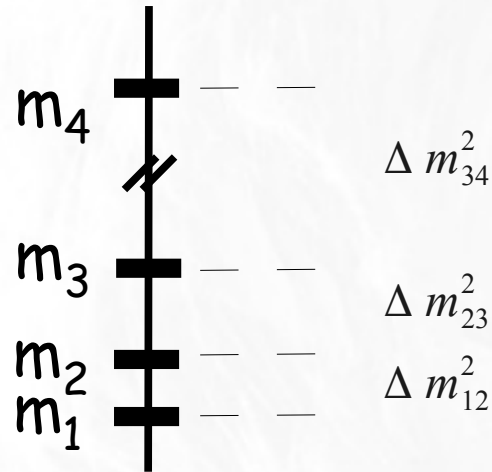
$\bar{\nu}_e$ in the LSND preferred
region but antinu results
consistent with LSND



explanation in terms of
sterile neutrinos

3+1 scheme

These states are
considered as
“degenerate”



m_4 is at a much higher scale,
around 1 eV^2 :
effective description in
terms of two-flavor

$$P_{\nu_{\alpha} \rightarrow \nu_{\beta}}^{(-)} = \delta_{\alpha\beta} - 4|U_{\alpha 4}|^2 (\delta_{\alpha\beta} - |U_{\beta 4}|^2) \sin^2 \left(\frac{\Delta m_{41}^2 L}{4E} \right)$$

$$\sin^2 2\vartheta_{e\mu} = 4|U_{e4}|^2 |U_{\mu 4}|^2 \text{ of } \nu_{\mu}^{(-)} \rightarrow \nu_e^{(-)} \text{ transitions}$$

$$\sin^2 2\vartheta_{ee} = 4|U_{e4}|^2 (1 - |U_{e4}|^2) \text{ of } \nu_e^{(-)} \text{ disappearance}$$

$$\sin^2 2\vartheta_{\mu\mu} = 4|U_{\mu 4}|^2 (1 - |U_{\mu 4}|^2) \text{ of } \nu_{\mu}^{(-)} \text{ disappearance}$$

ν_e appearance

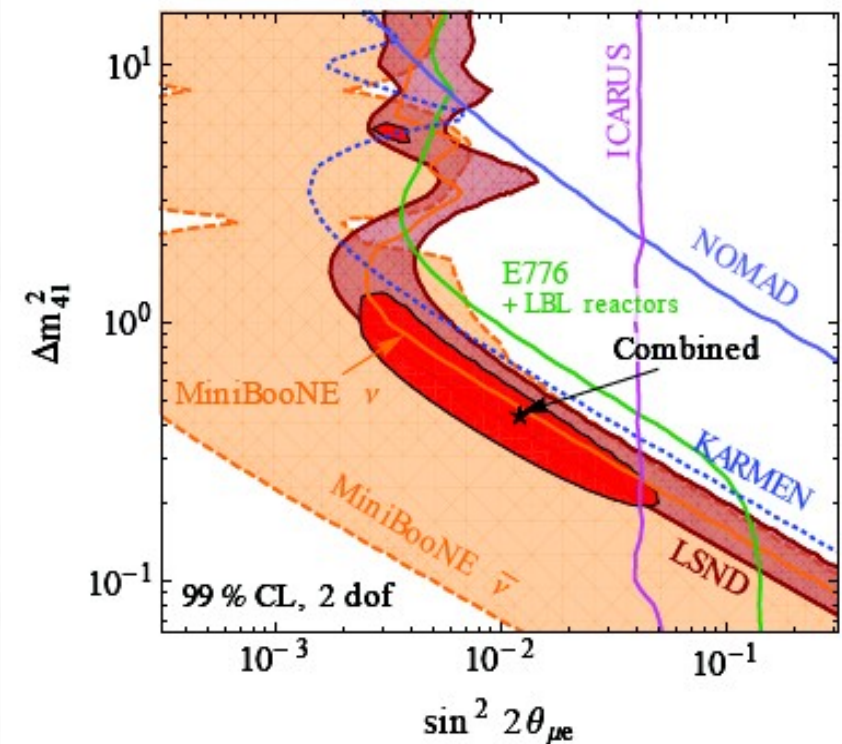
Global fit of ν_e appearance data
are consistent

$$\sin^2 2\theta_{\mu e} = 0.013$$

$$\Delta m_{41}^2 = 0.42 \text{ eV}^2$$

$$\chi^2_{\min}/\text{dof} = 87.9/66$$

Kopp, Machado, Maltoni,
Schwetz2013



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ν_e disappearance

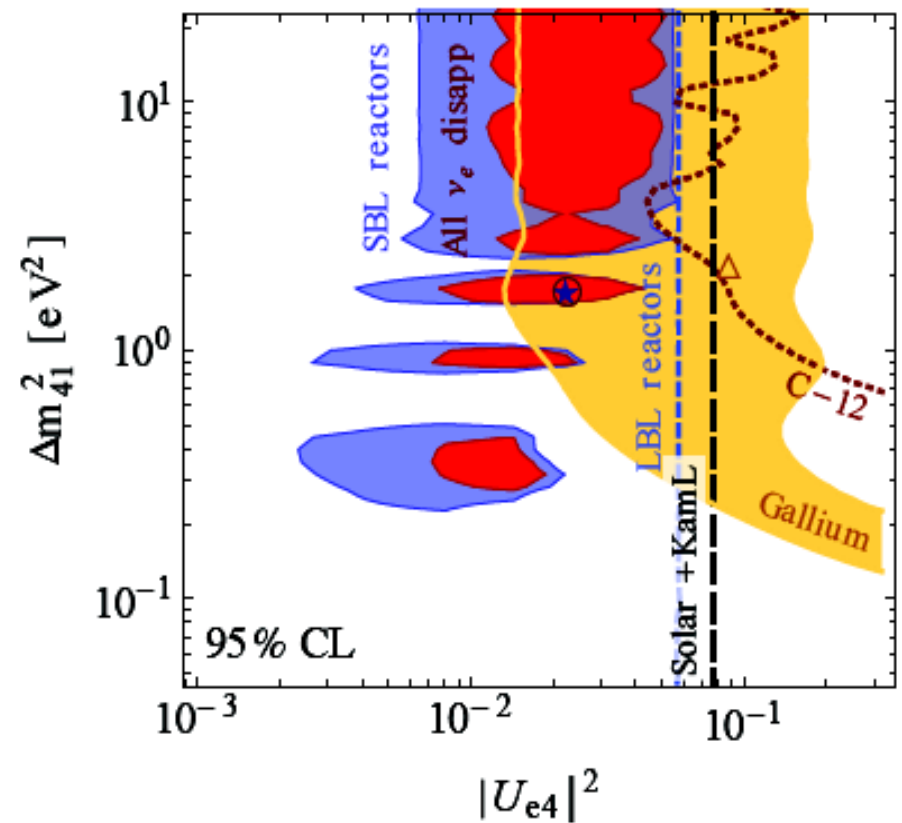
- Global fit on ν_e disappearance data are **consistent** among themselves

Kopp, Machado, Maltoni, Schwetz2013

$$\sin^2 2\theta_{ee} = 0.09$$

$$\Delta m_{41}^2 = 1.78 \text{ eV}^2$$

$$\chi^2_{\min}/\text{dof} = 403/427$$



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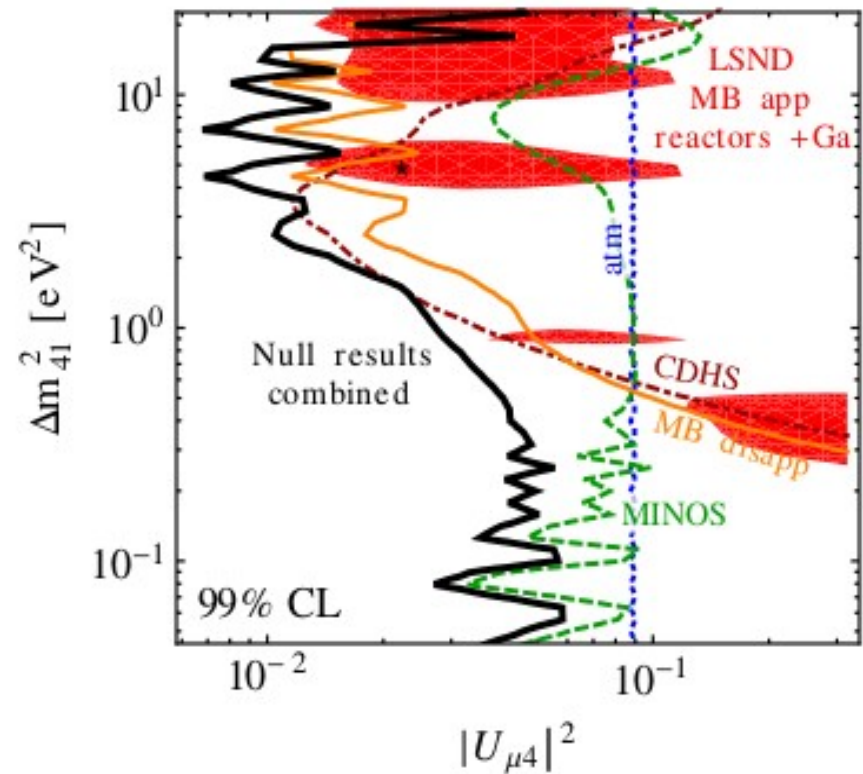
ν_μ disappearance

- Global fit on numu disappearance data:

Kopp, Machado, Maltoni,
Schwetz2013



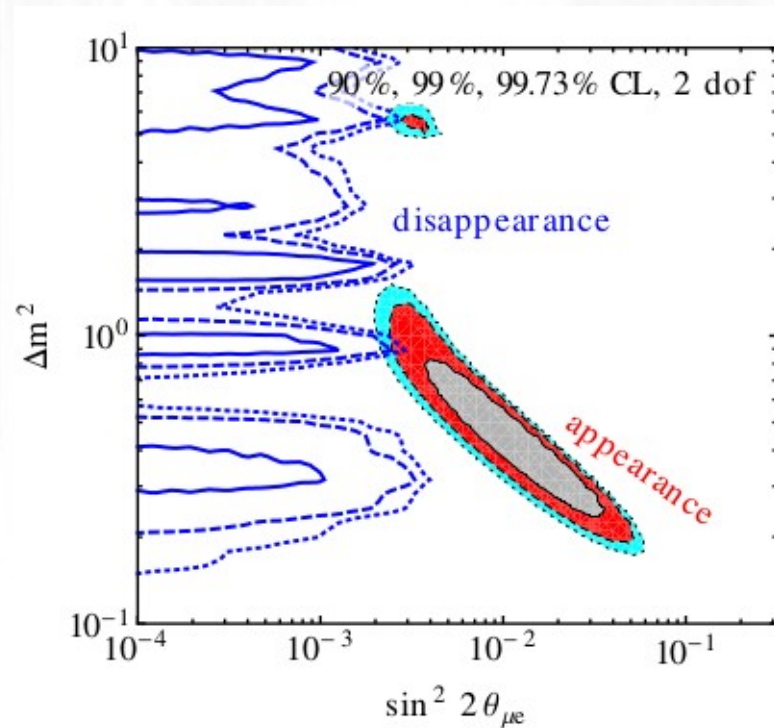
no signal \rightarrow strong constraints
on masses and mixing



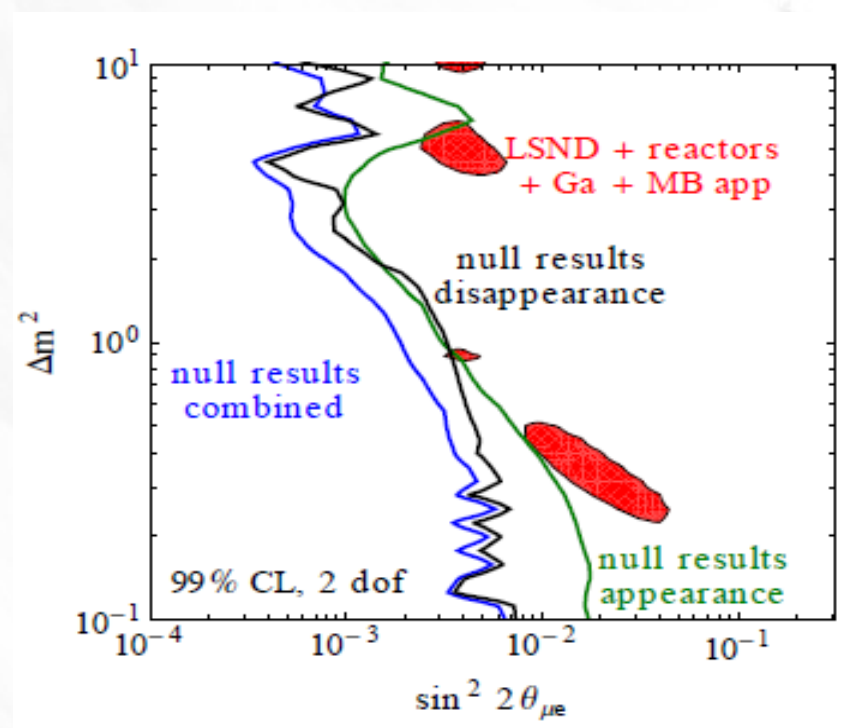
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Global picture

Tension between appearance
and disappearance



Tension between exp's with
and without signal



The (hard) job of a theorist

Take hints from experiments seriously
And explain:

Values of the
mixing angles



Smallness of masses

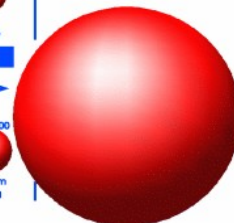
- Some ideas...

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On the ν masses

LEPTONS		
Electron Neutrino Mass -0	Muon Neutrino -0	Tau Neutrino -0
Electron .511	Muon 105.7	Tau 1.777
QUARKS		
Up Mass: 5	Charm 1.500	Top ~160,000
Down 8	Strange 160	Bottom 4.250



Easy part: neutrino Yukawa couplings smaller than those of the other fermions
Neutrinos are Dirac fermions: we have to introduce a right-handed neutrino field

neutrinos: $Y_\nu \bar{\psi}_L \tilde{H} \nu_R$

electrons: $Y_e \bar{\psi}_L H e^c$

$$\frac{Y_\nu}{Y_e} \sim 10^{-5}$$

But we want to go beyond
this "unnatural" scheme...

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Neutrino mass terms

we assume the existence of ν_L and the SM singlet ν_R

must be conserved: $|\Delta I| = 0$

Weak isospin	ν_L	ν_R	$H = (h^+, h^0)$
I	$1/2$	0	$1/2$
I_3	$1/2$	0	$(+1/2, -1/2)$

	ν	$\bar{\nu}$
Lepton number	1	-1

- Dirac mass term

(same for quarks and leptons)

lepton number L is conserved

$$L_D = m_D \bar{\psi}_L \tilde{H} \nu_R$$

- Majorana mass term

lepton number L is **not** conserved

$$L_M = m_M \nu_R^T \nu_R$$

The see-saw mechanism

- Total lagrangian

$$L_m = m_D \bar{\psi}_L \tilde{H} \nu_R + m_M \nu_R^T \nu_R$$

Electroweak symmetry breaking \rightarrow see-saw

$$\langle H \rangle = \begin{pmatrix} 0 \\ v \end{pmatrix} \rightarrow m_\nu = \begin{pmatrix} 0 & m_D \\ m_D^T & m_M \end{pmatrix}$$

$$m_\nu = -m_D^T \frac{1}{m_M} m_D$$

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The see-saw mechanism

- An indicative numerical example



$$m_v \sim m_D^2 / m_M$$

$$\text{for } m_D \sim 100 \text{ GeV}, m_v \sim 0.05 \text{ eV}$$



$$m_M \sim 10^{14} - 10^{15} \text{ GeV}$$

Probe into GUT!

Mixing angles

Two different approaches and equally (not?) promising

- Models with non-trivial dynamics: means that the structure of the mixing matrix is determined by discrete symmetries

such symmetries are motivated by the fact that the data themselves suggest rotations with fixed special angles ($\frac{1}{2}$, $1/3$...)

- permutational groups like A_4 , S_4 ...

- Models where the main idea is that there is no need of introducing additional symmetries to explain the mixing angles

In such models, the *chance* plays the fundamental role (anarchical models and variants)

Special mixing matrices

- mixing angles are obtained from the diagonalization of the mass matrix

$$m_{\nu}^{Diag} = U^T m_{\nu} U$$

- Good starting point suggested by the data:

$$\sin^2 \theta_{23} = \frac{1}{2} \quad \sin^2 \theta_{13} = 0$$

$$\sin^2 \theta_{12} = \frac{1}{3} \quad (TBM)$$

Tri-Bimaximal Mixing

$$\sin^2 \theta_{12} = \frac{1}{2} \quad (BM)$$

Bimaximal Mixing

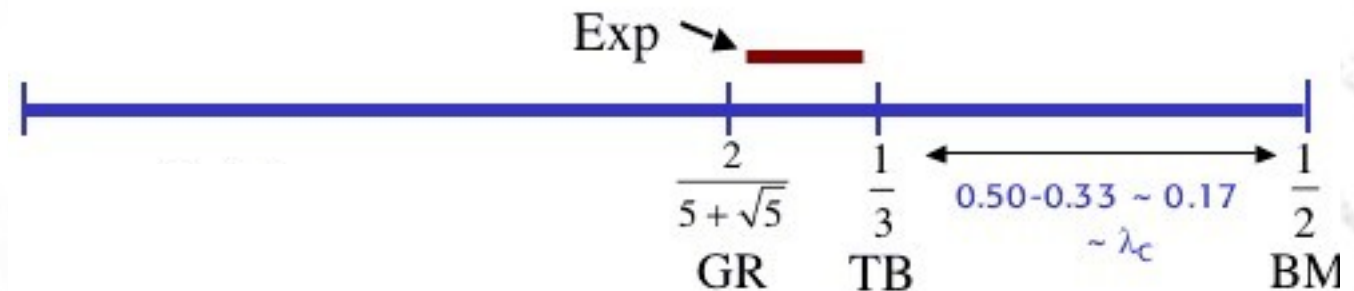
$$\sin^2 \theta_{12} = \frac{2}{5 + \sqrt{5}} \quad (GR)$$

Golden Ratio

Special mixing matrices

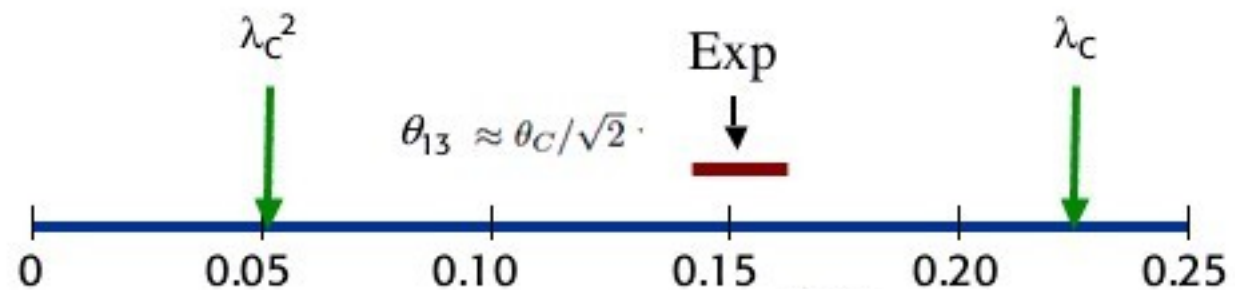
- How good are such starting points ?

$$\sin^2 \theta_{12}$$



$\lambda_c =$ Cabibbo angle

$$\sin \theta_{13}$$



Corrections are needed to stay on the experimental data

Special mixing matrices

- in models with no baroque dynamics, all mixing angles receive corrections of the same order of magnitude

TBM

$$\sin^2 \theta_{12} = \frac{1}{3} + O(\lambda_C^2) \quad \sin^2 \theta_{23} = \frac{1}{2} + O(\lambda_C^2) \quad \sin \theta_{13} = O(\lambda_C^2)$$

ok

ok

not really good

BM

$$\sin^2 \theta_{12} = \frac{1}{2} + O(\lambda_C) \quad \sin^2 \theta_{23} = \frac{1}{2} + O(\lambda_C) \quad \sin \theta_{13} = O(\lambda_C)$$

ok

ok

ok

This pattern seems to be favored

Special mixing matrices

Possible origin of corrections

- U_{PMNS} receives contributions from the charged lepton diagonalization

$$\nu_\alpha = U_{\alpha i}^\nu \nu_i$$

diagonalizes the neutrino mass matrix

$$l_\alpha = U_{\alpha i}^l l_i$$

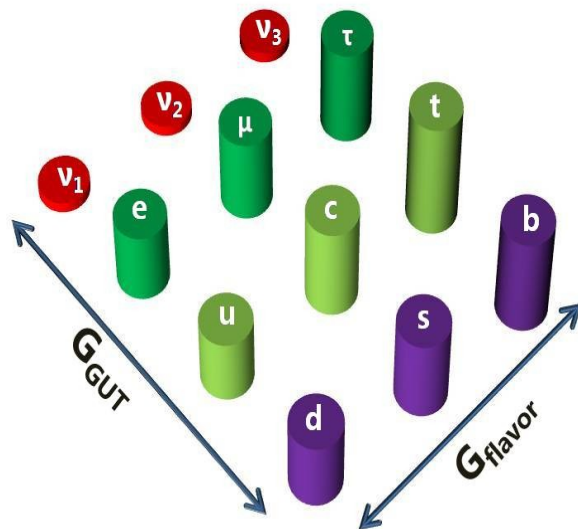
diagonalizes the charged lepton mass matrix

$$\bar{l}_\alpha \gamma_\mu \nu_\alpha W^\mu \rightarrow \underbrace{U_{\alpha i}^{+l} U_{\alpha j}^\nu}_{U_{\text{PMNS}}} \bar{l}_i \gamma_\mu \nu_j W^\mu$$

Charged current

Additional symmetries

- The previous patterns are easily obtained using *flavor symmetries*



- **gauge** symmetries act on members of particle multiplets
- **flavor** symmetries act on different families

Vantages: strong correlation among the entries of the mass matrices, so less free parameters \rightarrow predictability

Additional symmetries at work

- The models work as follows:

Altarelli-Feruglio2012

Theory invariant
under G_F

Flavor group

- Symmetry breaking of the flavor group: new scalar fields Φ in the theory with non vanishing vevs

Residual symmetry

given by a subset of the generators of G_F

in the neutrino sector $G_\nu \rightarrow U_\nu$

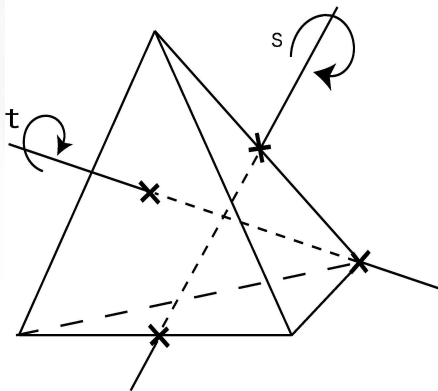
in the charged lepton sector $G_l \rightarrow U_l$

$$U_{PMNS} = U_l^\dagger U_\nu$$

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A possible flavor group: A_4

A_4 : is the group of even permutation of 4 objects
(also the symmetry of a tetrahedron)



generators of the group

$$T = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \omega^2 & 0 \\ 0 & 0 & \omega \end{pmatrix}, \quad S = \frac{1}{3} \begin{pmatrix} -1 & 2 & 2 \\ 2 & -1 & 2 \\ 2 & 2 & -1 \end{pmatrix}$$

The 12 elements are obtained considering all possible even permutations of 1234. They belong to **4 conjugacy classes**... given a of G , $\{g^{-1}ag, \forall g \in G\}$

A_4 has 4 irreducible representations

- **three** singlets **1**, **1'** and **1''**
- **1** triplet **3**

A possible flavor group: A_4

After breaking of A_4

- charged lepton mass matrix (residual symmetry generated by T)

$$m_e^{(0)} = v_d \begin{pmatrix} y_e & 0 & 0 \\ 0 & y_\mu & 0 \\ 0 & 0 & y_\tau \end{pmatrix} \eta$$

$$U_l = I$$

- neutrino mass matrix (generated by a **non**-diagonal generator S of A_4)

$$m_\nu^{(0)} = \begin{pmatrix} x & y & y \\ y & x+z & y-z \\ y & y-z & x+z \end{pmatrix}$$

$$U_\nu = \begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & -\sqrt{\frac{1}{2}} \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \end{pmatrix}$$

$$\sin^2 \theta_{23} = \frac{1}{2} \quad \sin^2 \theta_{12} = \frac{1}{3} \quad \sin^2 \theta_{13} = 0$$

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A possible flavor group: A_4

After breaking of T and S

- Charged lepton rotation

$$U_e = \begin{pmatrix} 1 & c_{12}^e \xi & c_{13}^e \xi \\ -c_{12}^{e*} \xi & 1 & c_{23}^e \xi \\ -c_{13}^{e*} \xi & -c_{23}^{e*} \xi & 1 \end{pmatrix}$$

- neutrino rotation

$$U_\nu = \begin{pmatrix} 1 & c_{12}^\nu \xi' & c_{13}^\nu \xi' \\ -c_{12}^{\nu*} \xi' & 1 & c_{23}^\nu \xi' \\ -c_{13}^{\nu*} \xi' & -c_{23}^{\nu*} \xi' & 1 \end{pmatrix}$$

$$\sin^2 \theta_{23} = \frac{1}{2} + \mathcal{R}e(c_{23}^e) \xi + \frac{1}{\sqrt{3}} \left(\mathcal{R}e(c_{13}^\nu) - \sqrt{2} \mathcal{R}e(c_{23}^\nu) \right) \xi \longrightarrow \frac{\langle \Phi \rangle}{\Lambda} \sim O(0.1)$$

$$\sin^2 \theta_{12} = \frac{1}{3} - \frac{2}{3} \mathcal{R}e(c_{12}^e + c_{13}^e) \xi + \frac{2\sqrt{2}}{3} \mathcal{R}e(c_{12}^\nu) \xi$$

$$\sin \theta_{13} = \frac{1}{6} \left| 3\sqrt{2} (c_{12}^e - c_{13}^e) + 2\sqrt{3} (\sqrt{2} c_{13}^\nu + c_{23}^\nu) \right| \xi.$$

Altarelli, Feruglio, Merlo, Stamou '12

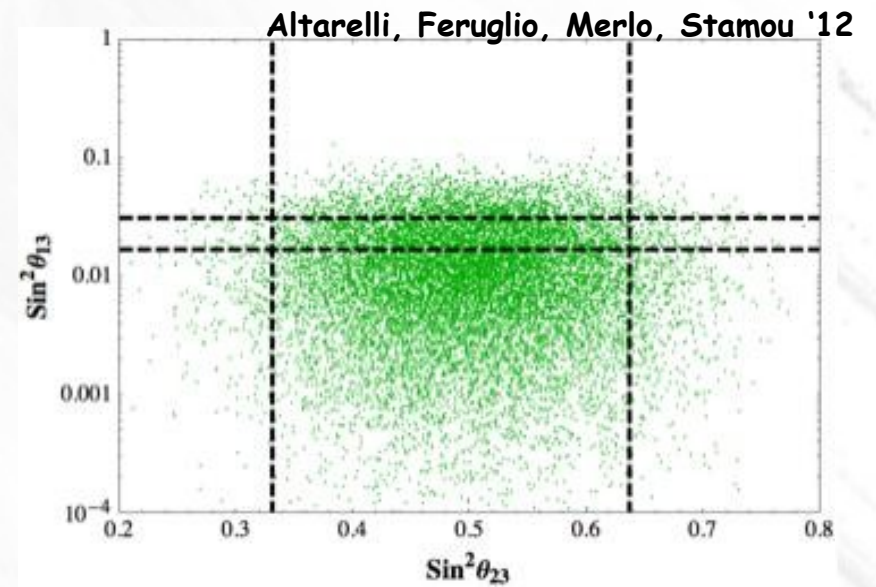
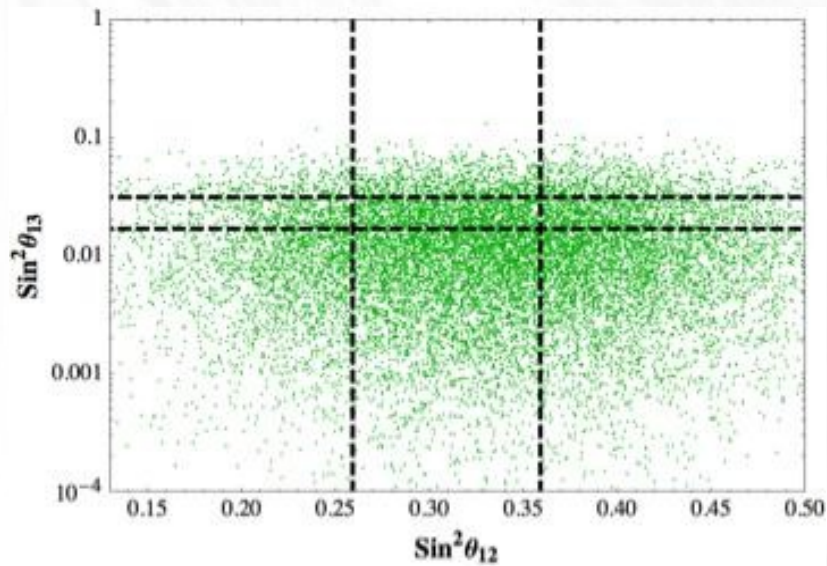
from charged lepton
rotation

from neutrino rotation

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Typical predictions of A4 models

c_{ij} = random complex with abs. value gaussian
around 1 with variance 0.5



Models with no special dynamics

- The **chance** is the basis of the success
- Only abelian U(1) to generate the hierarchies among fermions

- fields transform as: $\psi \rightarrow e^{i q_\psi} \psi$
- so a mass term transforms as:

$$y \bar{\psi}_L H \psi_R \rightarrow e^{i(-q_{\psi_R} + q_{\psi_L} + q_H)} y \bar{\psi}_L H \psi_R$$

If $(-q_{\psi_R} + q_{\psi_L} + q_H) = 0$ the mass term is allowed, otherwise we need a new scalar field θ with charge q_θ and vev v_θ :

$$y \bar{\psi}_L H \psi_R \left(\frac{\theta}{\Lambda} \right)^k \rightarrow e^{i(-q_{\psi_R} + q_{\psi_L} + q_H + k q_\theta)} y \left(\frac{v_\theta}{\Lambda} \right)^k \bar{\psi}_L H \psi_R$$

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 **Suppression factor**

A GUT example

- Standard Model particles in the 10 and $\bar{5}$ representations (3 copies)

$$\bar{5} = \begin{pmatrix} d_1^c \\ d_2^c \\ d_3^c \\ e \\ -\nu \end{pmatrix}_L, \quad 10 = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & u_3^c & -u_2^c & u_1 & d_1 \\ -u_3^c & 0 & u_1^c & u_2 & d_2 \\ u_2^c & -u_1^c & 0 & u_3 & d_3 \\ -u_1 & -u_2 & -u_3 & 0 & e^c \\ -d_1 & -d_2 & -d_3 & -e^c & 0 \end{pmatrix}_L$$

1 = right-handed neutrino

- SU(5) mass terms:

$$m_{up} \sim 10 \times 10$$
$$m_d = m_e^T \sim 10 \times \bar{5}$$

$$m_{\nu_D} \sim \bar{5} \times 1$$

$$m_M \sim 1 \times 1$$

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Models with no special dynamics

Choosing appropriate U(1) charges we can get several mass matrices structures:

- Anarchycal models (A)

$$\begin{aligned} q_{\bar{5}} &= (0,0,0) \\ q_{10} &= (3,2,0) \\ q_1 &= (0,0,0) \end{aligned} \quad m_l = \begin{pmatrix} \lambda^3 & \lambda^3 & \lambda^3 \\ \lambda^2 & \lambda^2 & \lambda^2 \\ 1 & 1 & 1 \end{pmatrix}, \quad m_\nu = \begin{pmatrix} 1 & 1 & 1 \\ 1 & 1 & 1 \\ 1 & 1 & 1 \end{pmatrix}$$

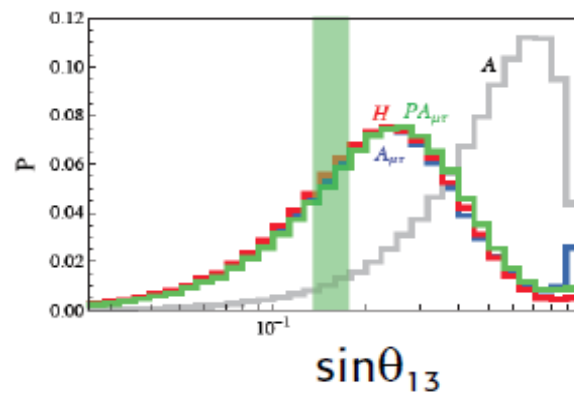
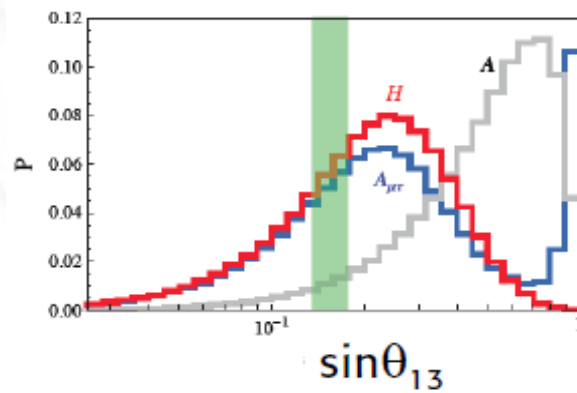
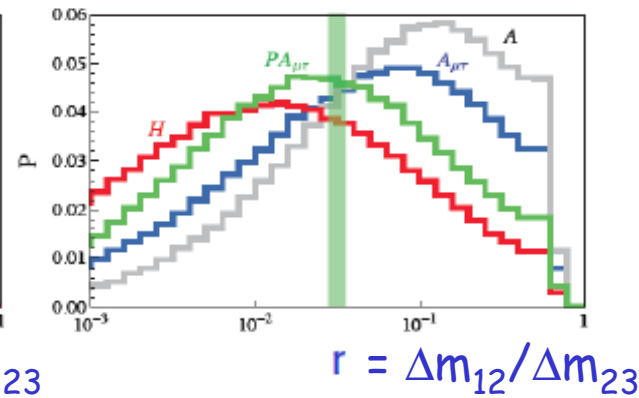
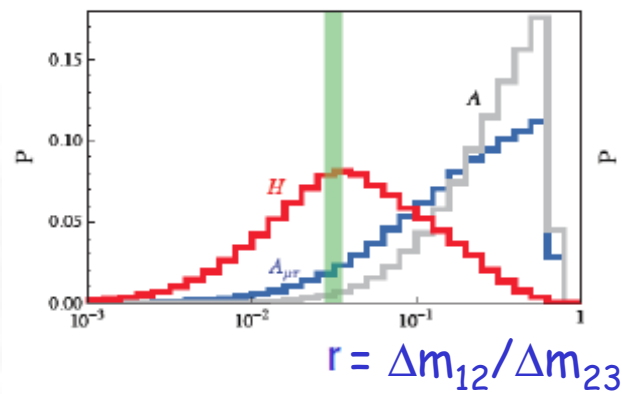
- Hierarchical model (H)

$$\begin{aligned} q_{\bar{5}} &= (2,1,0) \\ q_{10} &= (5,3,0) \\ q_1 &= (2,1,0) \end{aligned} \quad m_l = \begin{pmatrix} \lambda^7 & \lambda^6 & \lambda^5 \\ \lambda^5 & \lambda^4 & \lambda^3 \\ \lambda^2 & \lambda & 1 \end{pmatrix}, \quad m_\nu = \begin{pmatrix} \lambda^4 & \lambda^3 & \lambda^2 \\ \lambda^3 & \lambda^2 & \lambda \\ \lambda^2 & \lambda & 1 \end{pmatrix}$$

Models with no special dynamics

no see-saw

see-saw



message: H performs better than A

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The future (personal view)

Oscillation sector

- Better determination of the oscillation parameters and the mass pattern
- Check for new physics effects

Flavor sector

- Interplay of flavor symmetries and realistic GUT theories
- Differences among quarks and leptons

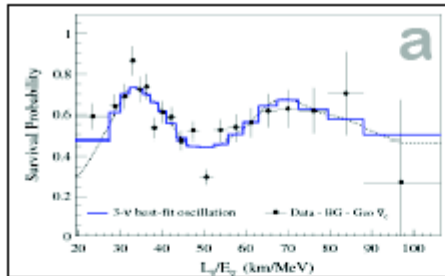
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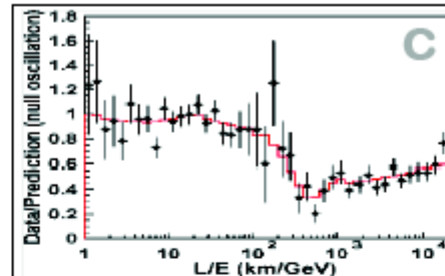
Global fit

Egidio Lisi, talk a
Moriond2015

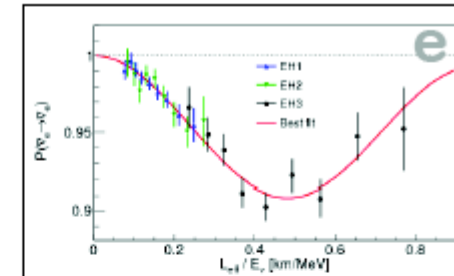
$e \rightarrow e$ (δm^2 , θ_{12})



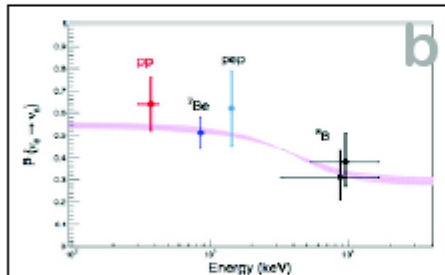
$\mu \rightarrow \mu$ (Δm^2 , θ_{23})



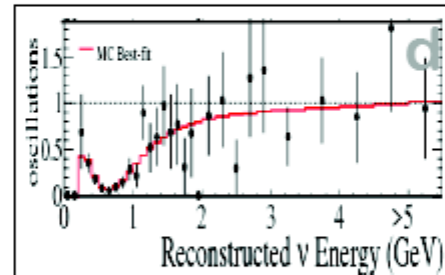
$e \rightarrow e$ (Δm^2 , θ_{13})



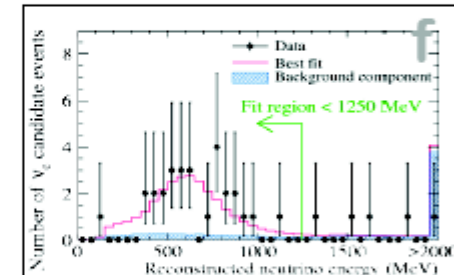
$e \rightarrow e$ (δm^2 , θ_{12})



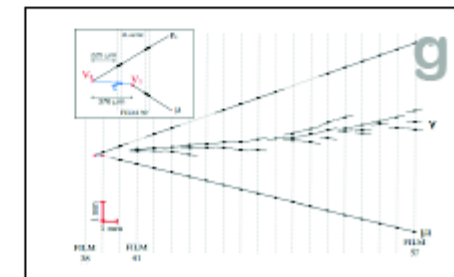
$\mu \rightarrow \mu$ (Δm^2 , θ_{23})



$\mu \rightarrow e$ (Δm^2 , θ_{13} , θ_{23})



$\mu \rightarrow \tau$ (Δm^2 , θ_{23})



Data from various types of neutrino experiments: (a) solar, (b) long-baseline reactor, (c) atmospheric, (d) long-baseline accelerator, (e) short-baseline reactor, (f,g) long baseline accelerator (and, in part, atmospheric).

(a) KamLAND [plot]; (b) Borexino [plot], Homestake, Super-K, SAGE, GALLEX/GNO, SNO; (c) Super-K atmosph. [plot], MACRO, MINOS etc.; (d) T2K [plot], MINOS, K2K; (e) Daya Bay [plot], RENO, Double Chooz; (f) T2K [plot], MINOS; (g) OPERA [plot], Super-K atmospheric.

D.Meloni

New results from Planck

For $T < m_e$, radiation content of the Universe is

$$\varepsilon_R = \varepsilon_\gamma + \varepsilon_\nu + \varepsilon_x$$

Ninetta Saviano,
talk at Moriond2015

non-electromagnetic contribution is parametrized in terms of effective neutrino species N_{eff}

$$\varepsilon_\nu + \varepsilon_x = \frac{7}{8} \frac{\pi^2}{15} T_\nu^4 N_{\text{eff}} = \frac{7}{8} \frac{\pi^2}{15} T_\nu^4 (N_{\text{eff}}^{\text{SM}} + \Delta N)$$

3.046

(relativistic degrees of freedom)

Extra radiation, for example from sterile neutrinos

Planck 2015:

$$N_{\text{eff}} = 3.15 \pm 0.46$$

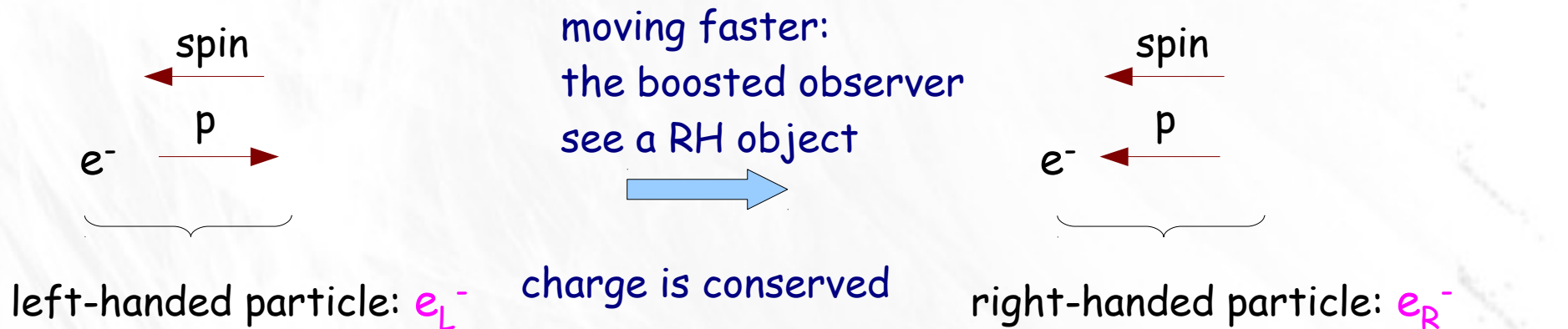
D.Meloni



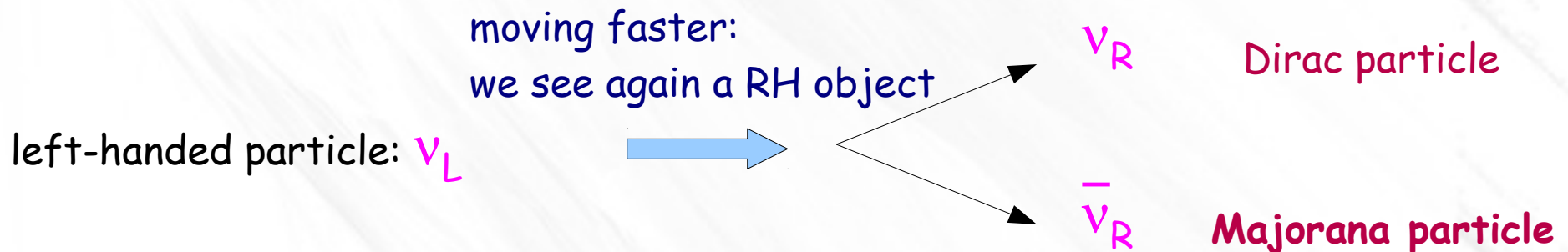
Not a large room for sterile states!

A peculiarity of the neutrino

- For electrons: 4 different helicity states and all of them are needed



- For neutrinos: from experiments we have identified ν_L and $\bar{\nu}_R$ only



if $\nu_L \sim \bar{\nu}_R$ no additive quantum numbers are conserved

A4

take this as an
example

- A_4 is the discrete group of even permutations of 4 objects ($4!/2 = 12$ elements) generated by S and T

$$S^2=T^3=(ST)^3=1$$

The action of the generators S and T can be assigned as follows:

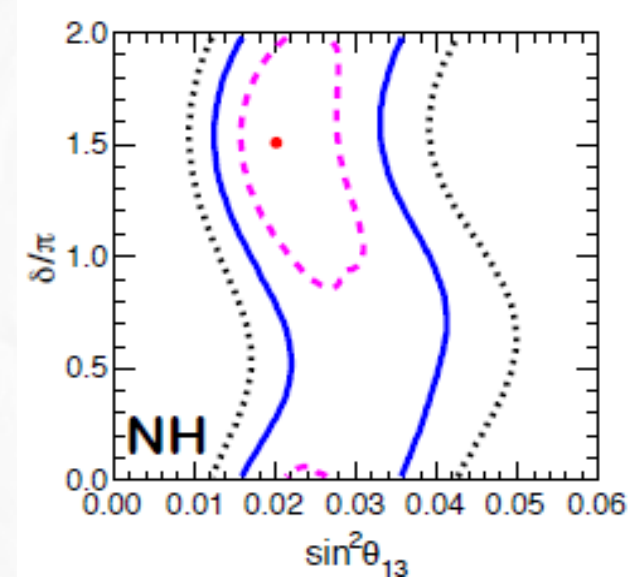
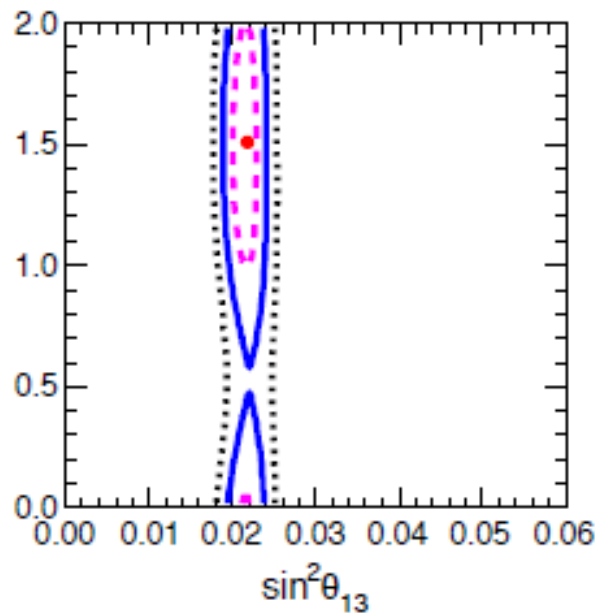
$$\textcolor{red}{S}: (1234) \rightarrow (4321) \quad \textcolor{red}{T}: (1234) \rightarrow (2314)$$

- irreducible representations:
a triplet and 3 different singlets $\textcolor{red}{3}, \textcolor{red}{1}, \textcolor{red}{1'}, \textcolor{red}{1''}$ (promising for 3 generations)
- invariance under S and T is automatic while A_{23} is not contained in A_4
(2-3 symmetry happens in A_4 if $1'$ and $1''$ symm. breaking flavons are absent or have equal VEV's)

A comment on the CP violating phase

Egidio Lisi,
Moriond2015

- Long Baseline experiments
(T2K) indicates $\delta \sim 3/2 \pi$

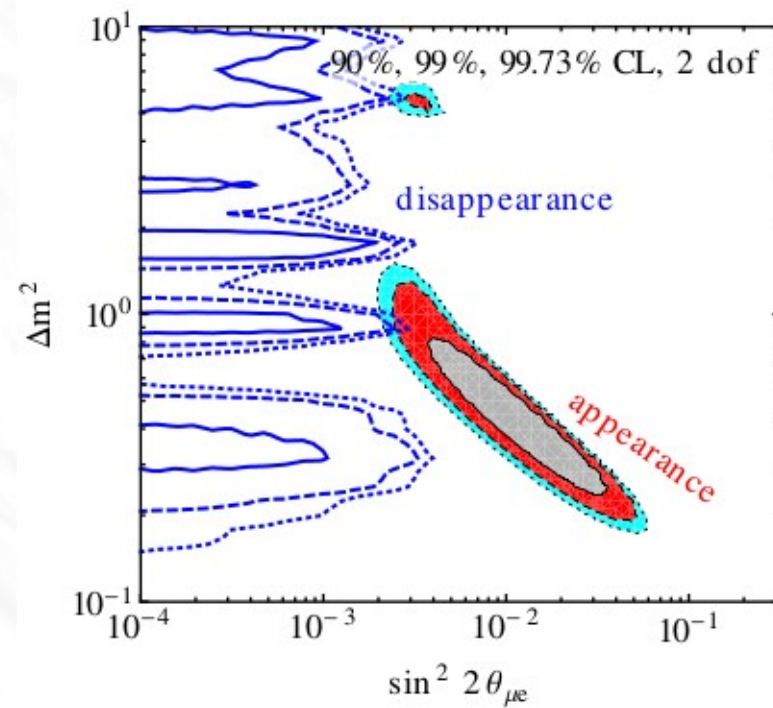


- Reactor experiments model
the CL form for $\sin^2 \theta_{13} \sim 0.02$

D. Meloni

Appearance-disappearance tension

$$\sin^2 2\theta_{\mu e} \approx \frac{1}{4} \sin^2 2\theta_{ee} \sin^2 2\theta_{\mu\mu}$$



D.Meloni